Outlier Detection in DESI Data

presentation for the LBNL HEP ML Group



Dark Energy Spectroscopic Instrument

U.S. Department of Energy Office of Science

 $0.7M \text{ Ly-} \alpha \text{ QSOs}$

1.7M QSOs

0.5 < z < 3.5

z = 3

18M ELGs

0.6 < z < 1.6

z = 2

(emission line galaxies):

z = 4

DESI: ~35,000,000 galaxies: 2021-2026 $r = 5.0 \, \text{Gpc/h}$ r = 4.0 Gpc/hTargeting several classes of astronomical objects r = 3.0 Gpc/hSample selection based on r = 2.0 Gpc/hbroadband photometry so samples are not pure 4M LRGs $r = 1.0 \frac{Gpc}{h}$ (luminous red galaxies): r 7 0.5 Gpc/h

0.4 < z < 1.0

10M bright galaxies z < 0.4

z = 1.5

z = 1

z = 0.7

z = 0.5

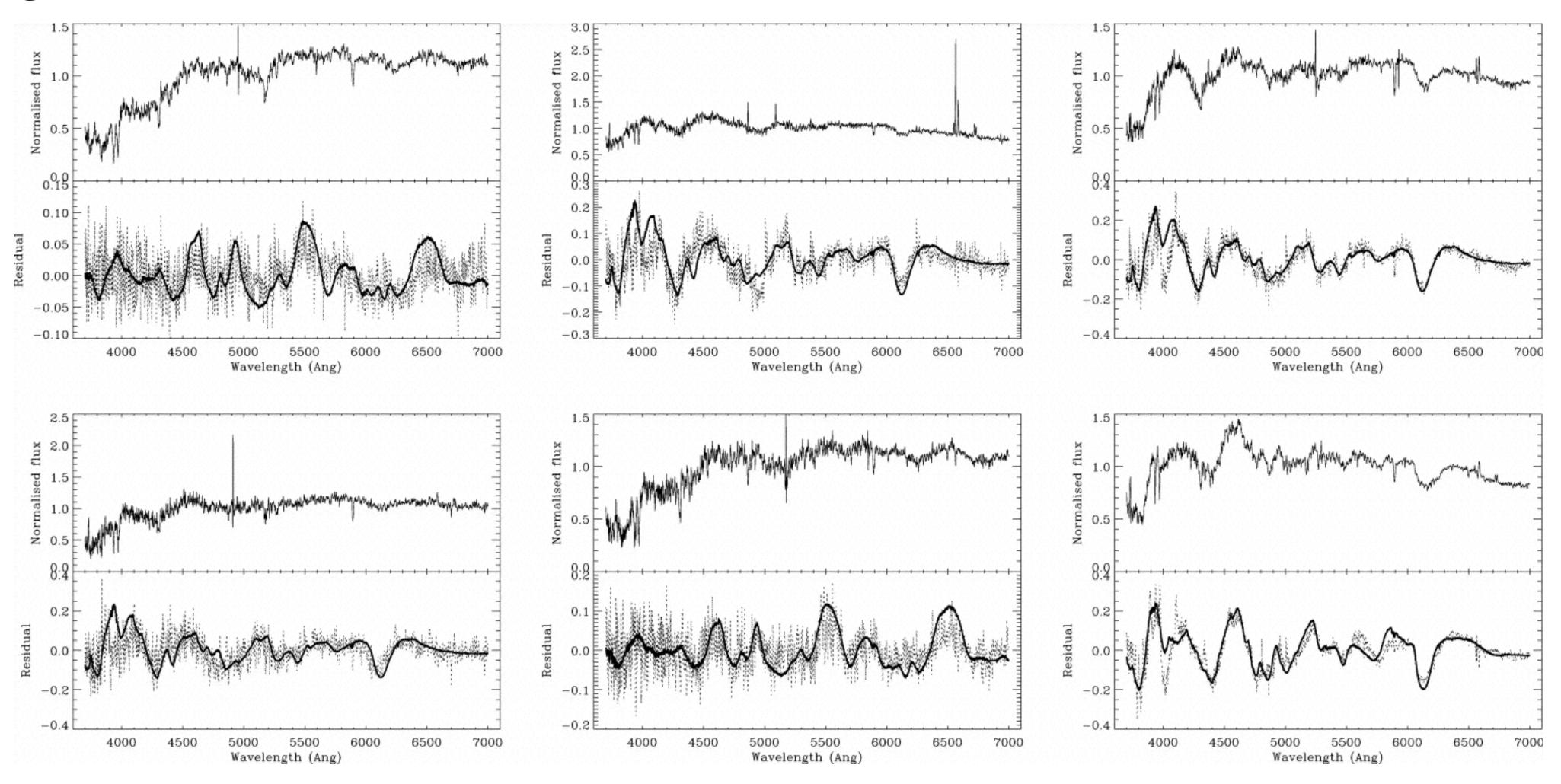
z=0.

MW Survey: ~ IOM stars

Observed Specta May Not Quite Look Like a Galaxy "Strange" Galaxy Spectra -> Transient Discovery in SDSS

Top:
Observed Spectrum

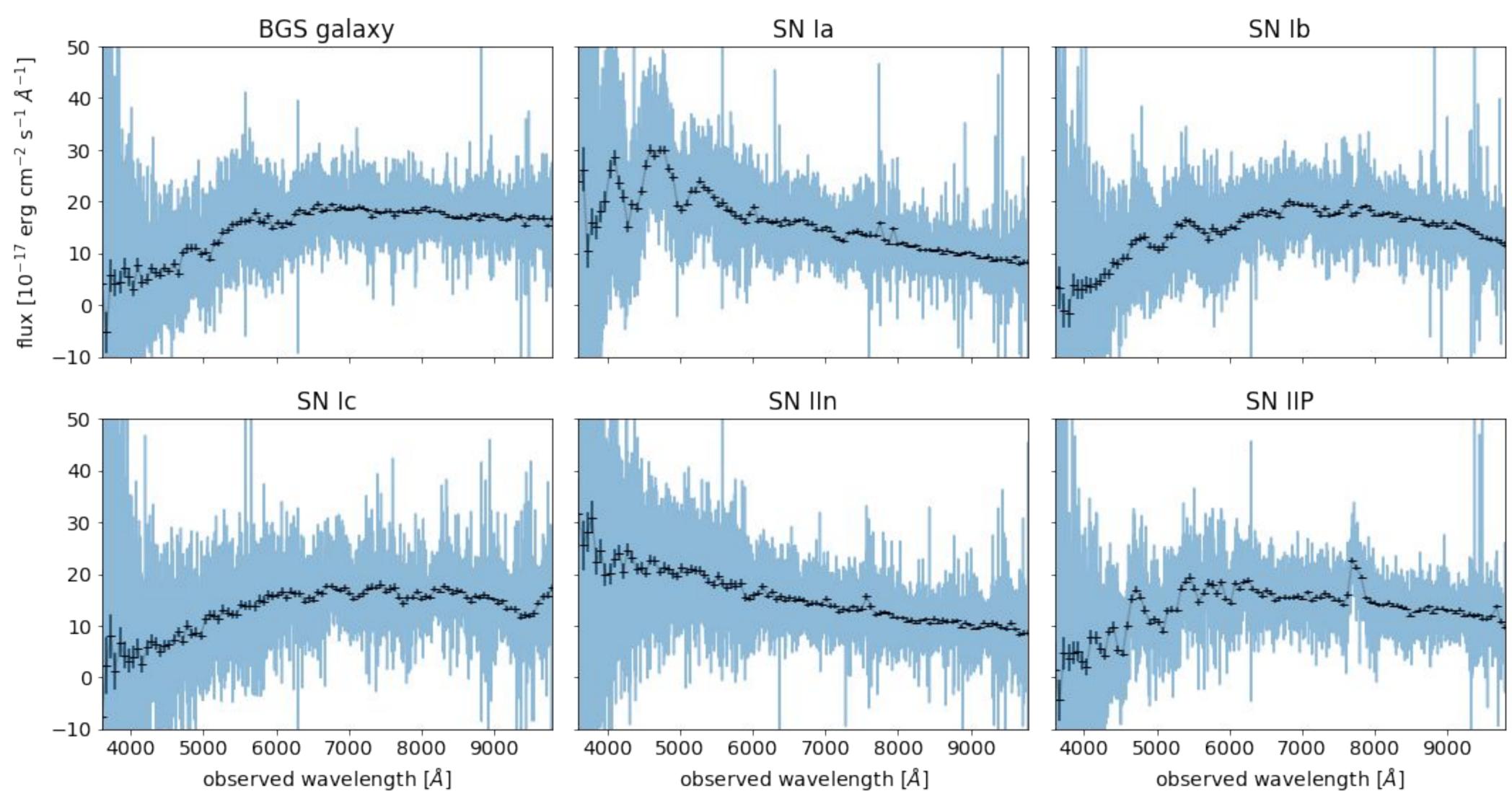
Bottom:
Observed Spectrum
- galaxy template
SN template





Simulated DESI Spectra

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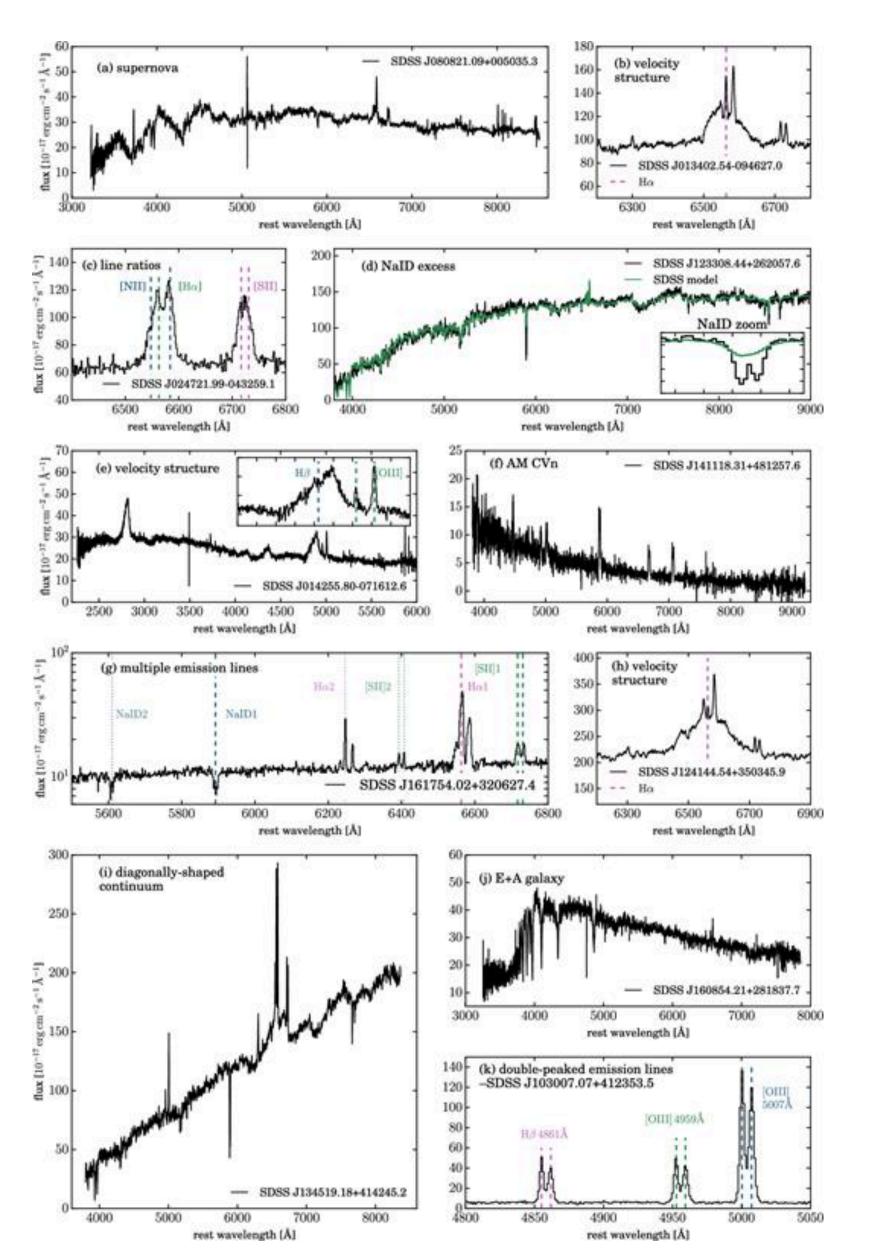


Data / Features

- Per exposure
 - Reduced Spectrum: Wavelength, Flux, Variance, Mask [array[7781]]
 - Metadata: Observing conditions, instrumental configuration
 - Postprocessing: Estimated redshift, ...
- Multiple exposures per object
- Pipeline version

Efforts Past and Present

- Sloan Digital Sky Survey (SDSS)
 - A large dataset in the can https://www.sdss.org/dr16/
 - Unsupervised Random Forest with 2M SDSS galaxies (see right)
- Efforts within DESI
 - Algorithms
 - Template-based searches
 - Supervised Learning based on empirical transient models
 - Application
 - Missing: Unsupervised learning because of a lack of person power
 - Mobilizing follow-up resources for live follow-up



Gallery of SDSS outliers from unsupervised learning Baron & Poznanski (2016)

DESI Discovery Potential from Outlier Detection

- ~2 SNe per 15,000 "Bright" Galaxies
 - Core SNe not efficiently found in imaging transient searches
- 1-8 Tidal Disruption Events per year
- Galaxy-galaxy gravitational lenses
- Galaxies with unusual gas kinematics
- Galaxy pairs
- ???

My "Machine Learning" Background

- I am not a Machine Learning person
- Past and current ML interests
 - Large number (>1000) parameter fits: Hamiltonian Monte Carlo, multi-probe cosmology, differentiable universe
 - Non-analytic likelihoods: Approximate Bayesian Calculation
 - Non-linear dimension reduction: Diffusion maps
 - "Parameterless: Gaussian Processes
- Science interests that touch on ML
 - SN classification (LSST Brokers), SN la standardization, multi-probe cosmology analysis